

Properties of the Neutron

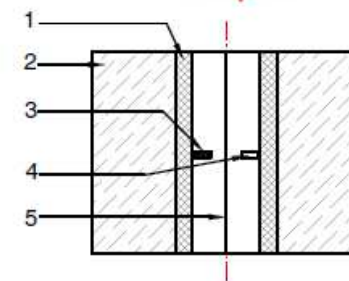
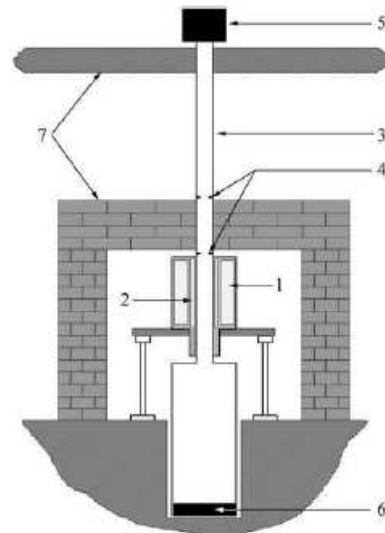
- Properties
 - Neutron in strong interactions
 - Electromagnetic interactions
 - Weak interaction
 - Neutron and Gravity
 - Other forces

Neutron in strong interaction

- Neutron-neutron scattering length

$$a_{nn}^{\text{str}} = -18.9 \pm 0.4 \text{ fm}, a_{pp}^{\text{str}} = -17.3 \pm 0.4 \text{ fm}$$

Recent idea: Pulsed reactor YAGUAR in Snezhinsk, Russia



[Furman et al., JPG 28, 2627 (2002);
Muzichka et al., NPA 789, 30 (2007)]

V INT, Seattle, WA, 3/25/20 - p.5/27

- Neutron-nucleus interactions

Electromagnetic interaction (I)

- Neutral particle
PDG value -0.2 ± 0.8
($10^{-21} e$)

- Radius of the neutron
(measurement of the EM FF)

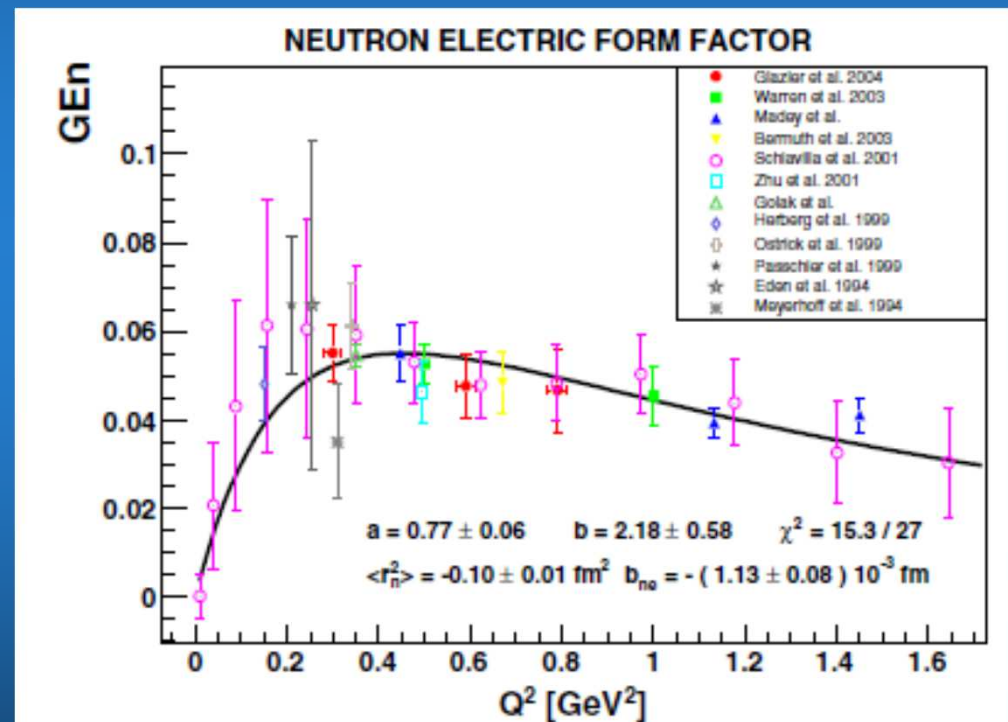


FIG. 4 (color online). The neutron form factor $G_E(\mathbf{q}^2)$ as a function of the momentum transferred \mathbf{q}^2 . The experimental data are taken from [29]; the solid curve is a two parameter fit using formula (19).

Electromagnetic interaction (II)

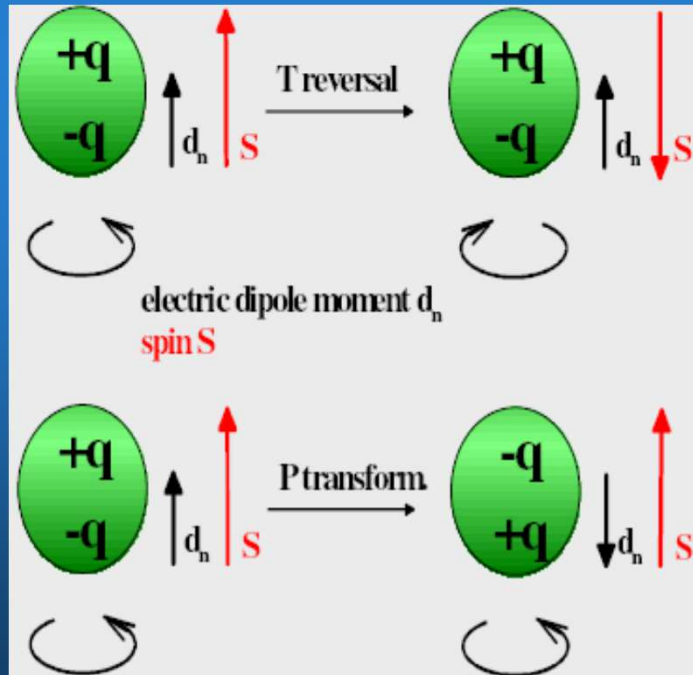
- Magnetic moment

PDG value $-1.91304272 \pm 0.00000045$ in nm

- Electric dipole moment (EDM)
- ...

Neutron Electric Dipole Moment (nEDM) – search for CP violation

$$\vec{d}_n = e \cdot \vec{r} = d_n \hat{s}$$

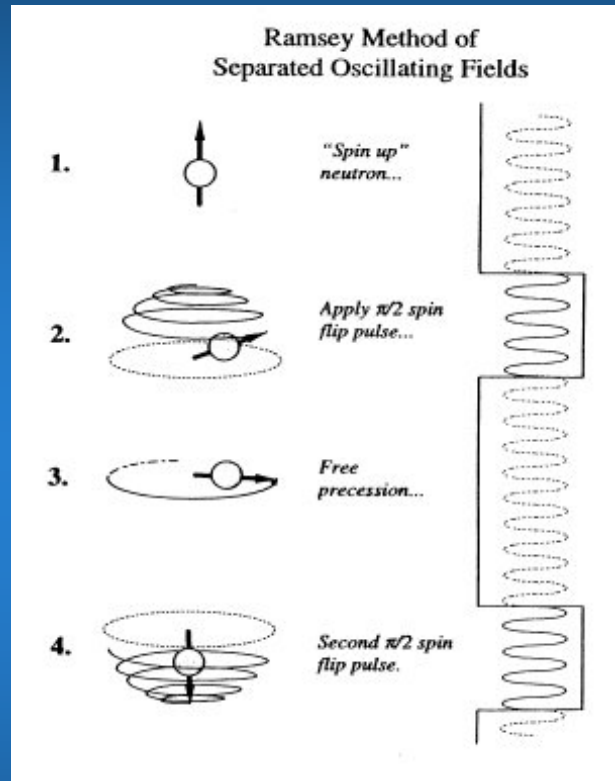


CP violation was observed only in the systems of neutral kaons and B mesons

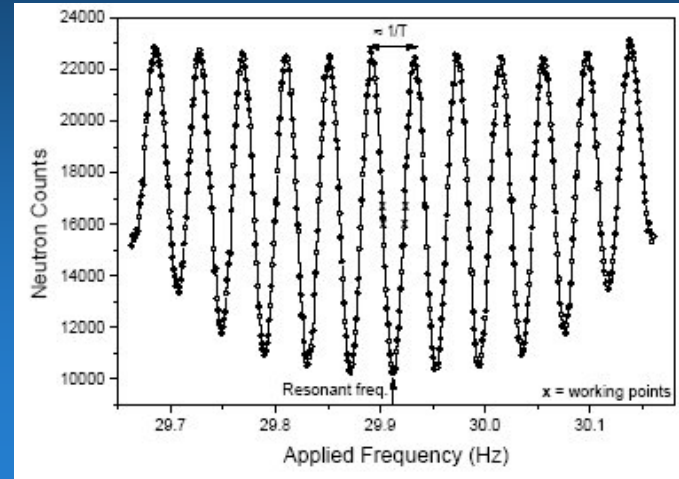
Existence of the $nEDM \neq 0$ means violation of **P** and of **T**

CPT Theorem \longrightarrow CP violation

Ramsey's Method

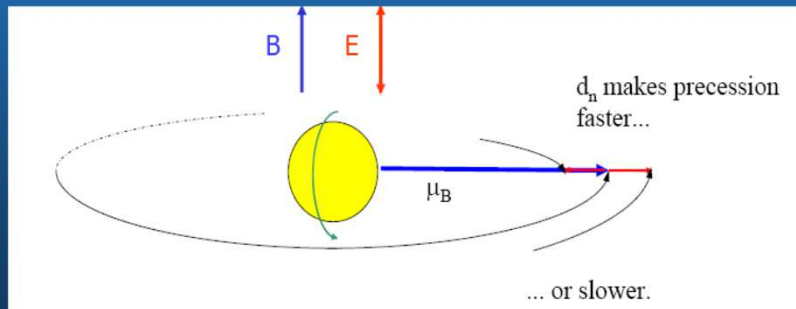
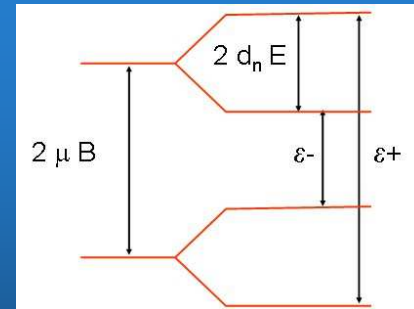


$N \uparrow$



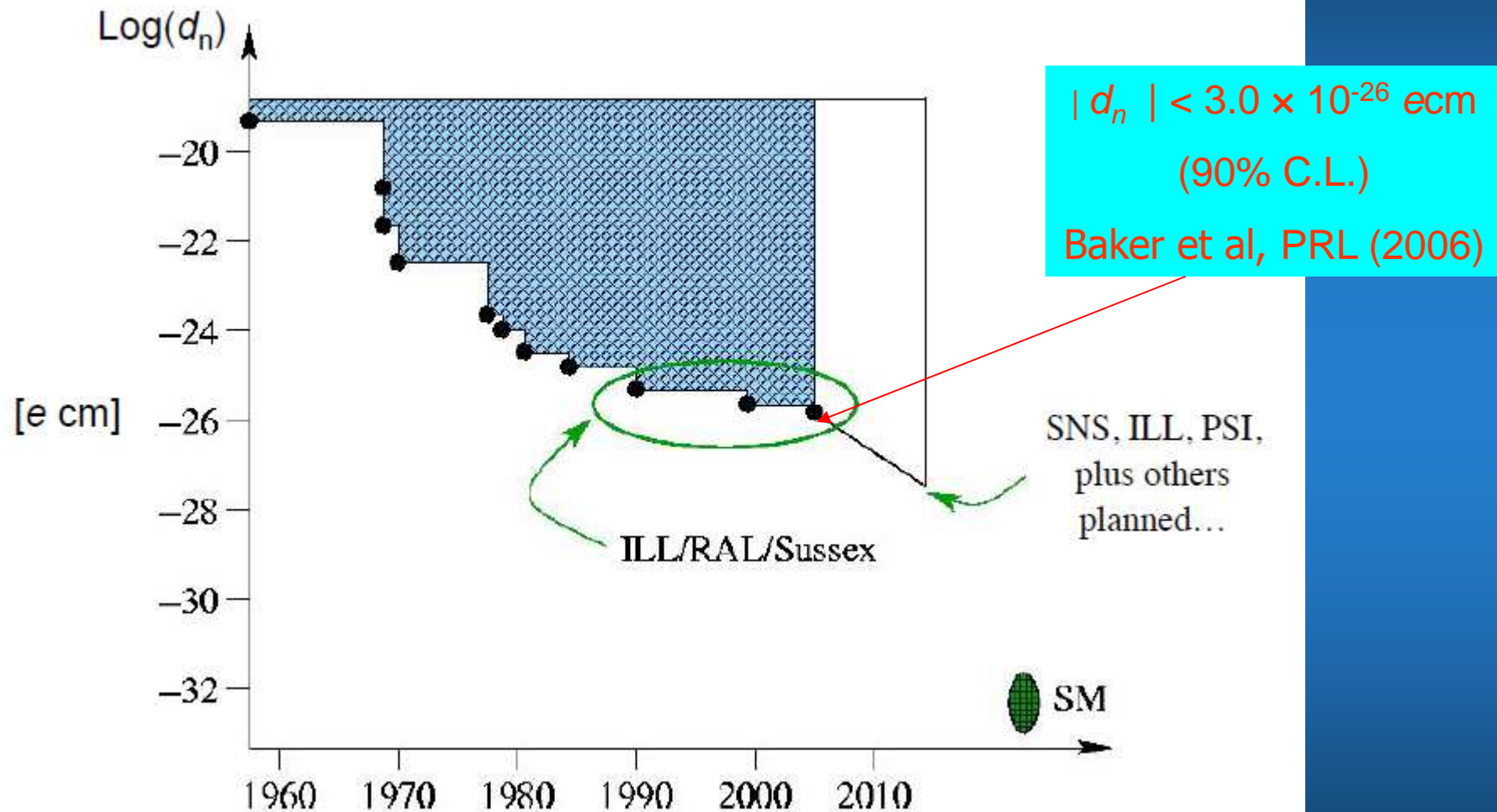
T

$$H = -2(\mu_n \vec{I} \cdot \vec{B} + d_n \vec{I} \cdot \vec{E})$$



$$\Delta\nu = 4d_n E / h$$

Progress in the neutron EDM bound



"It is fair to say that the neutron EDM has ruled out more theories (put forward to explain K_0 decay) than any experiment in the history of physics" R. Golub

Weak interaction

- Neutron lifetime
- Asymmetry coefficients in neutron decay

$$dW(\sigma, \mathbf{p}_e, \mathbf{p}_{\bar{\nu}}) \propto F(E_e) d\Omega_e d\Omega_{\bar{\nu}} \left\{ 1 + a \frac{\mathbf{p}_e \cdot \mathbf{p}_{\bar{\nu}}}{E_e E_{\bar{\nu}}} + \frac{bm_e}{E_e} + \langle \sigma_n \rangle \left(A \frac{\mathbf{p}_e}{E_e} + B \frac{\mathbf{p}_{\bar{\nu}}}{E_{\bar{\nu}}} + D \frac{\mathbf{p}_e \times \mathbf{p}_{\bar{\nu}}}{E_e E_{\bar{\nu}}} + R \frac{\boldsymbol{\sigma}_e \times \mathbf{p}_e}{E_e} + \dots \right) \right\}$$

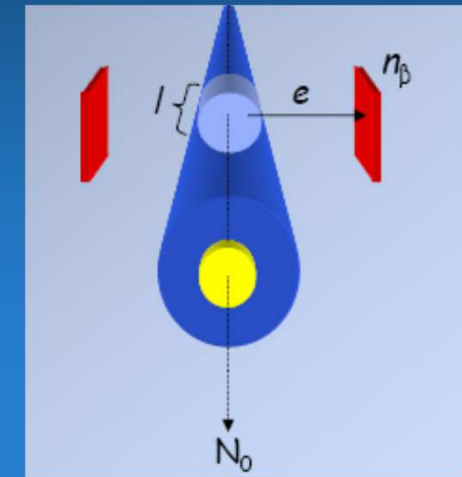
$$\lambda = G_A/G_V$$

$$a = \frac{1 - |\lambda|^2}{1 + 3|\lambda|^2}, \quad A = -2 \frac{|\lambda|^2 + \text{Re}(\lambda)}{1 + 3|\lambda|^2}, \quad B = 2 \frac{|\lambda|^2 - \text{Re}(\lambda)}{1 + 3|\lambda|^2}$$

- In flight experiment (CN) :

- To measure the neutron beam radioactivity:

$$n_{\beta} = \frac{dN}{dt} = -\frac{N_0}{\tau_n} e^{-\frac{l}{v \cdot \tau_n}}$$

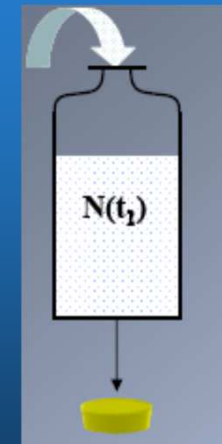


- Two measurements of absolute values

- Storage experiment (UCN) :

- To measure directly the decreasing of number of stored neutrons

$$\frac{1}{\tau_m} = \frac{1}{t_2 - t_1} \cdot \ln \frac{N(t_1)}{N(t_2)}$$



- Two relative measurements, but: $\tau_m^{-1} = \tau_n^{-1} + \tau_{loss}^{-1}$

Neutron lifetime

- **Standard Model**

$$|V_{ud}|^2 = \frac{(4908 \pm 4)s}{\tau(1 + 3\lambda^2)}$$

$$\lambda = \frac{g_A}{g_V}$$

$$\begin{pmatrix} d_w \\ s_w \\ b_w \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

- **Astrophysics**

- Solar cycle
- neutron stars formation
- Big Bang nucleosynthesis
- ...

« Typical » experiment with UCN (MamBo I)

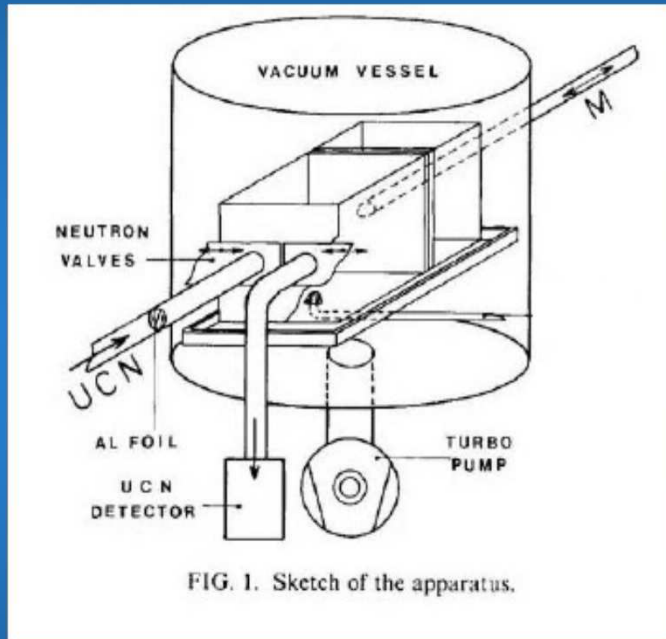


FIG. 1. Sketch of the apparatus.

$$\tau_{\beta} = (887.6 \pm 3) \text{ s}$$

Neutron Lifetime Measured with Stored Ultracold Neutrons
W. Mampe et al, Phys. Rev. Letters **63** (1989) 6

Idea :

- Storage volume (V) and its surface (S) are variables
- To measure the storage time as a function of the ratio V/S
- Extrapolate to infinite volume

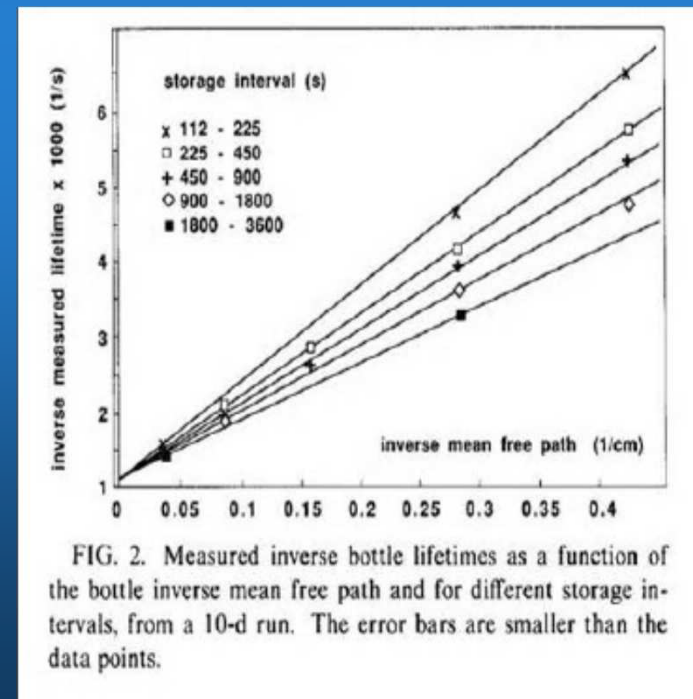


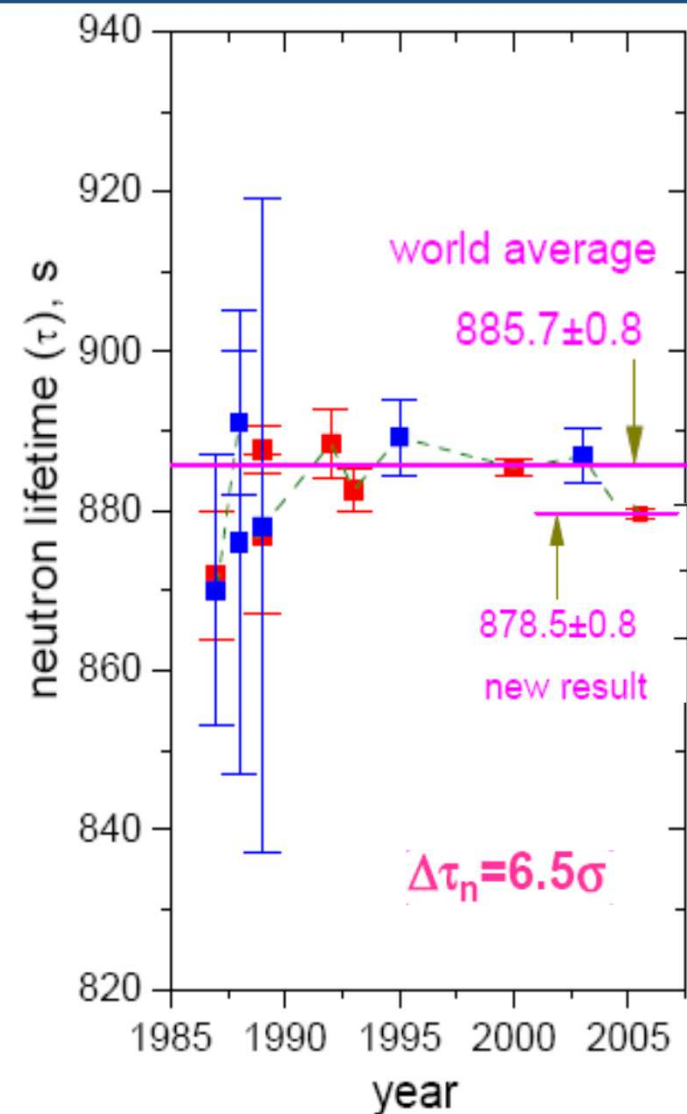
FIG. 2. Measured inverse bottle lifetimes as a function of the bottle inverse mean free path and for different storage intervals, from a 10-d run. The error bars are smaller than the data points.

$$\tau_n = (887.7 \pm 1.2 \text{ [stat]} \pm 1.9 \text{ [syst]}) \text{ s.}$$

Neutron lifetime

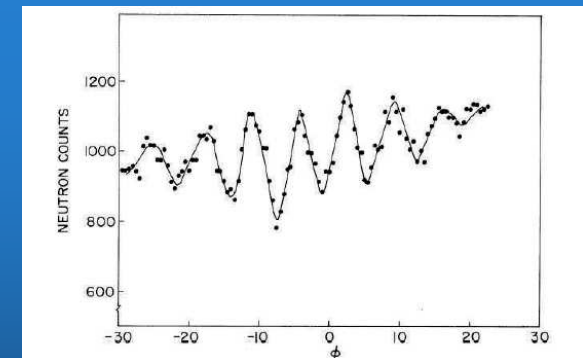
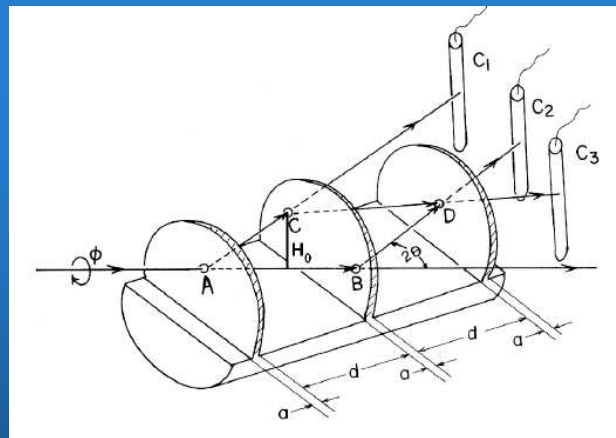
PRL (2013)

Lifetime τ [s]	Method	Ref./Year
878.5 ± 0.8	Storage of ultra-cold neutrons	A. Serebrov et al. 2005
886.8 ± 3.42	Neutron beam experiment	M.S. Dewey et al. 2003
885.4 ± 0.95	Storage of ultra-cold neutrons	S. Arzumanov et al. 2000
889.2 ± 4.8	Neutron beam experiment	J. Byrne et al. 1995
882.6 ± 2.7	Storage of ultra-cold neutrons	W. Mampe et al. 1993
888.4 ± 3.1 ± 1.1	Storage of ultra-cold neutrons	V. Nesvizhevski et al. 1992
878 ± 27 ± 14	Neutron beam experiment	R. Kosakowski 1989
887.6 ± 3.0	Storage of ultra-cold neutrons	W. Mampe et al. 1989
877 ± 10	Storage of ultra-cold neutrons	W. Paul et al. 1989
876 ± 10 ± 19	Neutron beam experiment	J. Last et al. 1988
891 ± 9	Neutron beam experiment	P. Spivac et al. 1988
872 ± 8	Storage of ultra-cold neutrons	A. Serebrov et al. 1987
870 ± 17	Neutron beam experiment	M. Arnold et al. 1987
903 ± 13	Storage of ultra-cold neutrons	Y.Y. Kosvintsev et al. 1986
875 ± 95	Storage of ultra-cold neutrons	Y.Y. Kosvintsev et al. 1980
937 ± 18	Neutron beam experiment	J. Byrne et al. 1980
881 ± 8	Neutron beam experiment	L. Bondarenko et al. 1978
918 ± 14	Neutron beam experiment	C.J. Christensen et al. 1972
885.8 ± 0.9	world average 1998	H. Abele 2000



Neutron and Gravity

- (Used in the UCN production)
- Interference experiments (m_i vs m_g)
Collela, Overhauser, and Werner (1975)



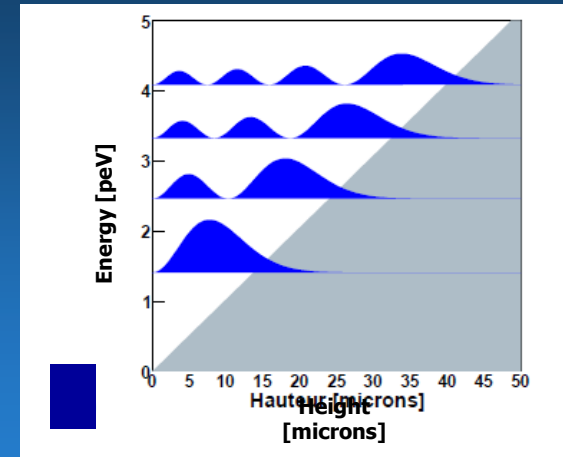
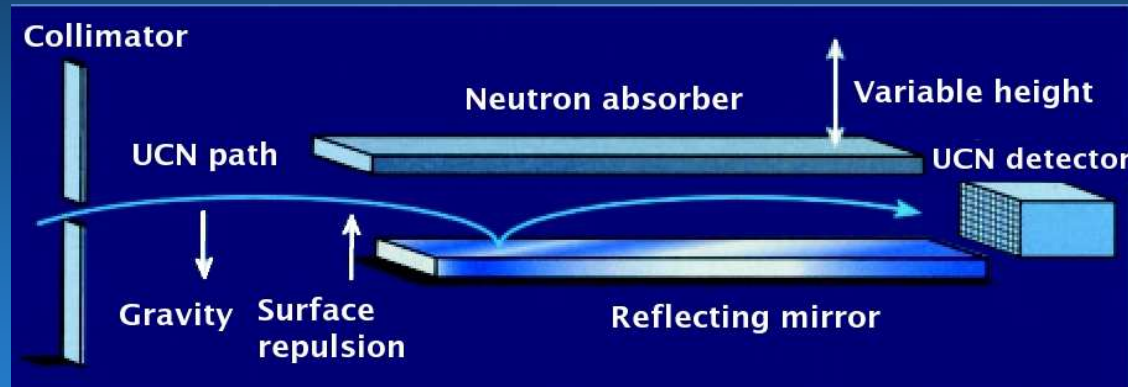
- Quantum states in gravity field

"Let us consider another possibility, an atom held together by gravity alone. For example, we might have two neutrons in a bound state. When we calculate the Bohr radius of such an atom, we find that it would be 10^8 light years, and that the atomic binding energy would be 10^{-70} Rydbergs. **There is then little hope of ever observing gravitational effects on systems which are simple enough to be calculable in quantum mechanics.**"

Brian Hatfield, in "Feynman Lectures on Gravitation" ;
R.P. Feynman, F.B. Morinigo, W.G. Wagner, Ed. Brian Hatfield

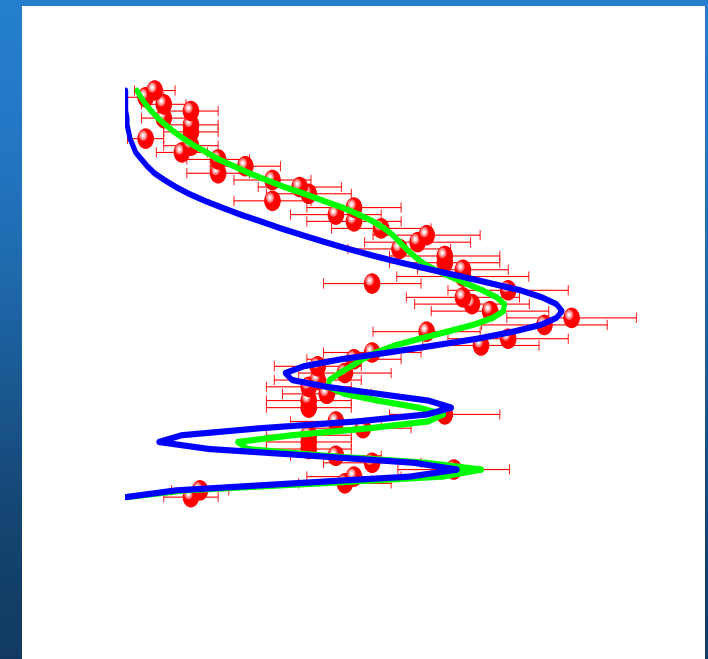
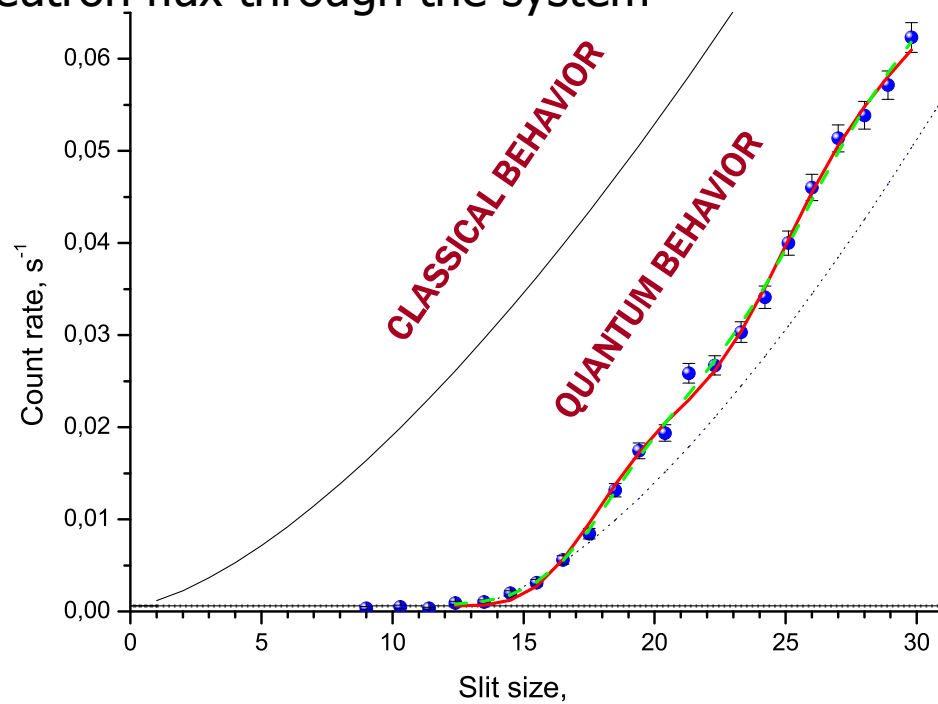
Addison-Wesley Publishing Company, **1995**, p. 11

Quantum System: Neutron in a Gravity Field



First « photo » of the wave function

Neutron flux through the system



Neutron in all interactions

- Strong interaction
- Weak interaction
- Electromagnetic interaction
- Gravitational interaction

A black rectangular box containing the handwritten text "What else?" in a light brown or tan color, written in a cursive script.

- Search for other interactions

Search for “5th force”

Scattering amplitude of slow neutrons on atoms

$$f(\mathbf{q}) = f_{\text{nucl}}(\mathbf{q}) + f_{\text{ne}}(\mathbf{q}) + \dots + f_{\text{v}}(\mathbf{q})$$

constant for low energies

$$f_{\text{ne}}(\mathbf{q}) = -b_{\text{ne}} (Z - f(Z, \mathbf{q}))$$

$$b_{\text{ne}} = - \left. \frac{2 m}{B m_e} \frac{dG_E(\mathbf{q}^2)}{d\mathbf{q}^2} \right|_{\mathbf{q}^2=0}$$

with an unknown contribution induced by a hypothetical interaction

$$V(r) = A \frac{g^2}{4\pi} \frac{\hbar c}{r} e^{-r/\lambda}, \quad \rightarrow \quad f_{\text{v}}(q) = -A \frac{g^2}{4\pi} \frac{c}{\hbar} \frac{2m\lambda^2}{1 + (q\lambda)^2}$$

Some existing limits

Random potential model

Different methods to determine b

Estimation of b_{ne}

